The New Physics Case for Muon Beam-Dump Experiments

Rikab Gambhir

Email me questions at [rikab@mit.edu!](mailto:rikab@mit.edu) Based on [Cesarotti, **RG**, [2310.16110\]](https://arxiv.org/abs/2310.16110)

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Cosmological Horizon

Rikab Gambhir – DPF Pheno – 16 May 2024 A **muon collider** is an exciting future prospect for our community! 2020 2030 2040 2050 Average lifespan for American male born in 1998 2060 2070 Today \overline{N} **High Luminosity Large Hadron Collider (HL-LHC) LHC Cloud of Possible Futures Muon Collider Future Circular Collider** Why are muons so great? Ask me later! **DUNE**

Source: How are Muons Born?

Using PIP-II at Fermilab as a template…

Beam Dump

Beam Dump Fundamentals

Produce weakly coupled, long lived new particles with a large boost! Junk (residual muons, SM decays, etc) deflected and/or absorbed by shield

Nominal COM Energy: $\sqrt{s} = \sqrt{2E_0M}$ $l_{\rm NP} \approx \left(\frac{E_0}{\rm TeV}\right) \times \left(\frac{g}{10^{-6}}\right)^{-2} \times \left(\frac{m_{\rm NP}}{10 \text{ MeV}}\right)^{-2} \times 100 \text{m}$ Particles can travel a long distance:

Beam Dump Fundamentals

How many new physics particles do we expect to see? Depends on target, cross section, experiment geometry, and total number of muons:

$$
\frac{dN}{dx} = N_{\mu} \frac{N_0 \rho l_0}{A} \frac{d\sigma}{dx} \left(e^{\frac{L_{\text{tar}}}{l_{\text{NP}}}} - 1 \right) e^{-(L_{\text{tar}} + L_{\text{sh}})} \left(1 - e^{-L_{\text{dec}}/l_{\text{NP}}} \right)
$$

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New Physics – Bremsstrahlung Production

Given a beam dump, what kind of new physics ϕ can we search for?

$$
\mu(p) + N(P_i) \rightarrow \mu(p') + N(P_f) + \phi(k)
$$

Where ϕ is either spin-0 or spin-1 and either parity-even or parity-odd

Cross Sections – Calculation

Simplify the calculation … use Weizsäcker-Williams!

$$
\mu(p) + \gamma(q) \to \mu(p') + \phi(k)
$$

Approximation works best when the photon (virtuality *t*) is nearly on shell:

$$
t_{\rm min} \approx \left(\frac{m_\phi^2}{2E}\right)^2
$$

Sets the minimum beam energy we consider

Enhanced production when *t* is small – collinear emission, regulated by masses

$$
\left(\frac{d\sigma_{2\to 3}}{dx}\right)_{\text{IWW}} = \frac{g^2}{2\pi} \alpha \chi \frac{|\vec{k}|}{E} \frac{(1-x)}{x} \int_{-\infty}^{\tilde{u}_{\text{max}}} d\tilde{u} \frac{\mathcal{A}_{t=t_{\text{min}}}^{22}}{\tilde{u}^2} \text{ amplitude evaluated at } t_{\text{min}} \\ \text{integrated photon Flux} \qquad \qquad \uparrow \\ \chi \equiv \int_{t_{\text{min}}}^{t_{\text{max}}} dt \frac{t-t_{\text{min}}}{t^2} G(t) \text{ with the initial velocity of the system of the system.}
$$

Models	Today, I will only talk about these two.
Scalar	Seudoscalar
$\mathcal{L}_{int}^{S} \supset -ig_{S}\phi\bar{\psi}\psi$	Pseudoscalar
$e.g. \underline{\text{muonphilic}}$, leptophilic	
Vector	Axial Vector
$\mathcal{L}_{int}^{V} \supset -ig_{V}\psi_{\mu}\bar{\psi}\gamma^{\mu}\psi$	Axial Vector
$\mathcal{L}_{int}^{V} \supset -ig_{V}\psi_{\mu}\bar{\psi}\gamma^{\mu}\psi$	Axial Vector
$e.g. \underline{\text{Dark Photons}}, \, L_{\mu} - L_{\tau}$	Exalification

*ALPs are funny in our setup.

Models

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Scalar

$$
\mathcal{L}^{S}_{\rm int} \supset -ig_{S}\phi\bar{\psi}\psi
$$

e.g. muonphilic, leptophilic scalars

Vector
\n
$$
\mathcal{L}_{int}^{V} \supset -ig_{V}V_{\mu}\bar{\psi}\gamma^{\mu}\psi
$$
\ne.g. Dark Photons, L_{μ} - L_{τ}

Dark Photons

 P **Pseudoscalar**
Pseudoscalar Extend the SM with a broken *U*(1)' symmetry with a gauge boson *Z*'.

with parameter ε . Couples to SM charged particles via kinetic mixing with the photon

$$
\text{min}\{\text{min}\}
$$

$$
\mathcal{L} \supset \frac{1}{2} m_{Z'}^2 Z'^{\mu} Z'_{\mu} - \sum_{l \in e, \mu, \tau} i \epsilon e \left(\bar{l} \gamma^{\mu} l \right) Z'_{\mu}
$$

Classic DM candidate!

Dark photons are like light, except they're dark instead of light, and heavy instead of light

*ALPs are funny in our setup.

Results – Dark Photons

(a) Dark photon limits at $E_0 = m_h/2$.

(b) Dark photon limits at $E_0 = 5$ TeV.

Models

Scalar

$$
\mathcal{L}^{S}_{\rm int} \supset -ig_{S}\phi\bar{\psi}\psi
$$

e.g. muonphilic, leptophilic scalars

Vector $\mathcal{L}_{int}^{V} \supset -ig_{V}V_{\mu}\bar{\psi}\gamma^{\mu}\psi$ e.g. Dark Photons, *L - L*

Muonphilic

Model where the new scalar *only* couples to muons.

$$
\mathcal{L} \supset \frac{1}{2} m_{\phi}^2 \phi^2 - i \left[g_{\mu} \bar{\mu} \mu + g_{e} \bar{\ell} e + g_{\tau} \gamma \right] \phi
$$

5 operator: Can be generated by the dim

 $\phi LH\mu^c$

Can also arise as a Higgs-like coupling with $g_{\ell} \sim m_{\ell}/\Lambda$ with:

 $m_e \ll m_\mu$ and $m_\phi < m_\tau$

muonic couplings! **Muonic experiments probe**

See [[1902.07715\]](https://arxiv.org/abs/1902.07715) for example UV completions Can arise in e.g. Type III 2HDM

*ALPs are funny in our setup.

Results – Muonphilic

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Muon specific coupling, **no other limits** in this range!

Results – Muonphilic

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Muon specific coupling, **no other limits** in this range!

Results – Lots of models!

Results – Lots of models!

Considerations

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To obtain these results, we have made a lot of *conservative* simplifying assumptions – can do much better with a thorough experimental analysis!

- **Detector**: We assume a binary "Did you see a particle or not?". Much better sensitivity / background rejection achievable with tracking, calorimetry, timing … !
- **● Thin Targets**: We assumed the target is thin to simplify material calculations. This can be relaxed to increase the interaction volume!
- **Other visible signatures**: We only considered electrons, muons, and hadrons as visible final states. But other signatures (photons, MET, secondary decays) are possible!

A thorough analysis and simulation is required!

Considerations

Muon Beams are real!

 \sim 160 GeV, 10¹⁰ Muons

 $NAA\mu, \, \underline{2401.0170}$ [NA64μ, 2401.01708]

Earlier this year (Jan 2024)

Exploration of the Muon $q-2$ and Light Dark Matter explanations in NA64 with the

Conclusion

"This is our Muon Shot"

A **muon beam-dump experiment** provides an excellent opportunity for **new physics** on the road to a **muon collider**

New physics is possible *in the next few years* with as low as a **10 GeV** muon beam at a demonstrator facility! We don't need to wait.

Appendices

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Muon Stopping Power

Full Cross Section Formulas

$$
\left(\frac{d\sigma_{2\to 3}(p+P_i \to p'+k+P_f)}{d(p \cdot k)d(k \cdot P_i)}\right)_{WW} = \left(\frac{\alpha \chi}{\pi}\right) \left(\frac{E_0 x \beta_\phi}{1-x}\right) \times \left(\frac{d\sigma_{2\to 2}(p+q \to p'+k)}{d(p \cdot k)}\right)_{t=t_{\text{min}}} \left(\frac{d\sigma_{2\to 3}}{dx d\cos\theta}\right)_{WW} = \frac{g^2}{2\pi} \alpha |\vec{k}| E(1-x) \frac{\mathcal{A}_{t=t_{\text{min}}}^{22}}{\tilde{u}^2} \chi
$$

$$
\begin{split} &\mathcal{A}_{S,t=t_{\rm min}}^{2\to2} \approx \frac{x^2}{1-x} + 2 (m_\phi^2 - 4 m_\mu^2) \frac{\tilde{u} x + m_\mu^2 (1-x) + m_\mu^2 x^2}{\tilde{u}^2} \\ &\mathcal{A}_{P,t=t_{\rm min}}^{2\to2} \approx \frac{x^2}{1-x} + 2 m_a^2 \frac{\tilde{u} x + m_\mu^2 (1-x) + m_\mu^2 x^2}{\tilde{u}^2} \\ &\mathcal{A}_{V,t=t_{\rm min}}^{2\to2} \approx 2 \frac{2-2x+x^2}{1-x} + 4 (m_V^2 + 2 m_\mu^2) \frac{\tilde{u} x + m_\mu^2 (1-x) + m_\mu^2 x^2}{\tilde{u}^2} \\ &\mathcal{A}_{A,t=t_{\rm min}}^{2\to2} \approx \frac{4 m_\mu^2 x^2}{(m_A^2)(1-x)} + 2 \frac{2-2x+x^2}{1-x} + 4 (m_A^2 - 4 m_\mu^2) \frac{\tilde{u} x + m_\mu^2 (1-x) + m_\mu^2 x^2}{\tilde{u}^2} \end{split}
$$

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Beam Dump Fundamentals (Details)

How many new particles should we expect to see^{*?}

$$
\frac{dN}{dxdz} = N_{\mu} \frac{N_0 X_0}{A} \times \mathcal{BR} \times \int_{E_{\phi}}^{E_0} \frac{dE'}{E'} \int_0^T dt \ I(E'; E_0, t) \times E_0 \frac{d\sigma}{dx'} \bigg|_{x' \equiv E'/E_0} \frac{dP(z - \frac{X_0}{\rho}t)}{dz}
$$

*Assuming that the detector is wide enough to capture all emitted particles – we have chosen geometries and cutoffs such that this is approximately true. *Assuming 100% detection efficiency. *Assuming no SM backgrounds, taken care of by shields and/or absorption.

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Number of detected particles *N* as a function of their energy fraction *x* and decay length *z*

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Beam Dump Fundamentals (Muon Source)

How many new particles should we expect to see?

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\frac{dN}{dxdz} = N_{\mu} \frac{N_0 X_0}{A} \times \mathcal{BR} \times \int_{E_{\phi}}^{E_0} \frac{dE'}{E'} \int_0^T dt \ I(E'; E_0, t) \times E_0 \frac{d\sigma}{dx'} \Big|_{x' \equiv E'/E_0} \frac{dP(z - \frac{X_0}{\rho}t)}{dz}
$$
\nNumber of **muons on target**, 10¹⁸ - 10²²/ year

Beam Dump Fundamentals (Target)

How many new particles should we expect to see?

Beam Dump Fundamentals (Target)

How many new particles should we expect to see?

*N*₀: Avogadro's Number, ~10²³/mol *X0* : Material Decay Length *A*: Material atomic mass, ~10-100 g/mol

I(E'; E_q *, t)* = Radiative losses as the muon transverses a distance *t* through the target

Thin Target Approximation: No losses

$$
I(E'; E_0, t) = \delta(E' - E_0)
$$

Fixes the size of L_{tar} , otherwise requires more sophisticated material modeling …

Radiative losses for water and lead

Beam Dump Fundamentals (Decays)

How many new particles should we expect to see?

$$
\frac{dN}{dxdz} = N_{\mu} \frac{N_0 X_0}{A} \times \left| \mathcal{BR} \right| \times \int_{E_{\phi}}^{E_0} \frac{dE'}{E'} \int_0^T dt \ I(E'; E_0, t) \times E_0 \frac{d\sigma}{dx'} \Big|_{x' \equiv E'/E_0} \frac{dP(z - \frac{X_0}{\rho}t)}{dz}
$$

Branching ratio into *visible* **final** states (electrons, muons, hadrons)^{*}

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Probability of the new particle decaying at a (normalized) position *z*

$$
\frac{dP(l)}{dl} = \frac{1}{L_{\rm NP}} e^{-l/L_{\rm NP}}
$$

*Photons not considered, but definitely possible to do in other setups! *Hadrons obtained by using *R*-ratio measurements, OK below *Z* pole. * Tau decays are possible, but to be conservative we will assume we can't reconstruct them.

Beam Dump Fundamentals (Decays)

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$$

Branching ratio into *visible* **final** states (electrons, muons, hadrons)^{*}

Probability of the new particle decaying at a (normalized) position *z*

 $\Gamma_{\phi\to l^+l^-}=g_S^2{m_\phi\over 8\pi}\left(1-{4m_l^2\over m_\phi^2}\right)^{3/2}$ $\Gamma_{a\rightarrow l^+l^-}=g_P^2{m_\phi\over 8\pi}\left(1-{4m_l^2\over m_\phi^2}\right)^{1/2}$ $\Gamma_{V\to l^+l^-} = g_V^2 \frac{m_\phi}{12\pi} \left(1 + \frac{2m_l^2}{m_\phi^2} \right) \left(1 - \frac{4m_l^2}{m_\phi^2} \right)^{1/2}$ $\Gamma_{A\to l^+l^-} = g_A^2 \frac{m_\phi}{12\pi} \left(1 - \frac{4m_l^2}{m_\phi^2} \right)^{3/2}.$

Decay lengths and branching ratios depend on new particle mass, coupling, and spin/parity $\frac{dP(l)}{dl} = \frac{1}{|L_{\rm NP}} e^{-l/\text{L}_{\rm NP}}$

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Models – Cross Sections

Beam Dump Fundamentals

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Aside: PDFs, ISR, and FSR

Since particles radiate or can be non-perturbative, it is a lot easier to think about colliding **partons** instead.

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$$
\sigma \,\,=\,\, \sum_{a,b}\, \int\limits_0^1 {\rm d}x_a {\rm d}x_b \,\int\, f_{a/h_1}(x_a,\mu_F) f_{b/h_2}(x_b,\mu_F) \, {\rm d}\hat{\sigma}_{ab\to n}(\mu_F,\mu_R)
$$

 $f_a(x)$ = probability of finding parton *a* in the particle *h* with longitudinal momentum fraction *x*

Leptons are "cleaner" than hadrons because $f(x) \sim \Box(x-1)^x$

Useful Application: **Weizsäcker-Williams****

**Sometimes called the "Effective Photon Approximation" or "Effective Vector Approximation"

Timeline of Particle Physics

Why Muons?

Intersecting Storage Rings (ISR) CERN

"This is our Muon Shot"

Road to the Future $\frac{1}{2^{620}}$ $\frac{1}{2030}$ $\frac{1}{2040}$ $\frac{1}{2050}$

3. **"Electroweak** Boson Collider"

2. **Cleaner*** than *pp* collisions

Super Proton-Antiproton Synchrotron (SppS)

4. Direct probe of **2nd gen physics**

CERN 5. "It's just **fucking cool**" – Nima Arkani-Hamed, APS 2023

1. Achieve *much* **higher energies** than *e⁺ e -*

 Λ Department Λ Lots of R&D needed! A huge undertaking ...

Large Electron-Positron potential for discovering **new physics**! Interesting results for muon **This Talk**: R&D is worth it, because **muon test beams** already have the **beams as low as 10 GeV**!

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Large Hadron Collider (LHC)

Leptonic

Hadronic